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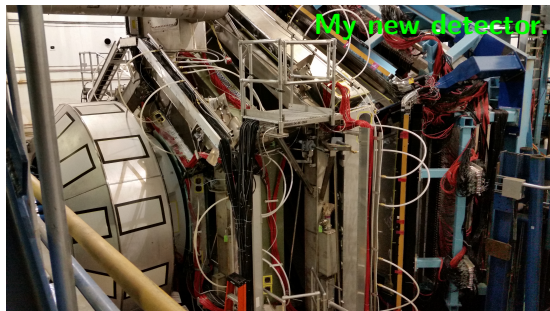
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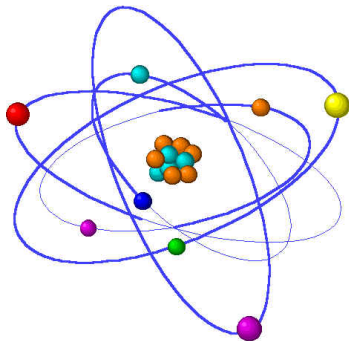
Albert Einstein

- “The more success the quantum theory has the sillier it looks.”

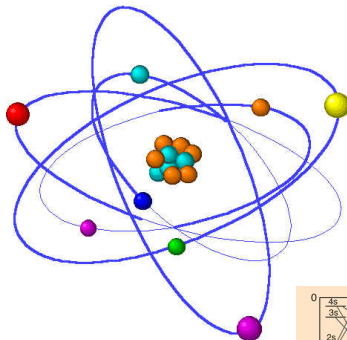
Albert Einstein

- 1 Start with a detector and take some data.
- 2 Develop the quantum program.
- 3 Apply the quantum program.
- 4 What are the classical alternatives?

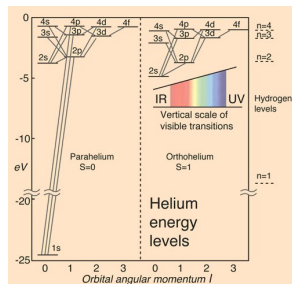
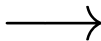
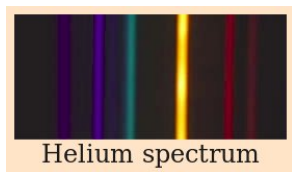




← A toy atom.
Sim is [here](#).

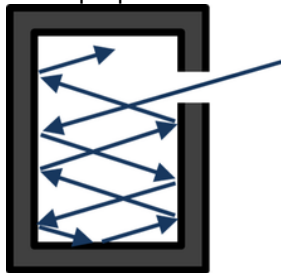


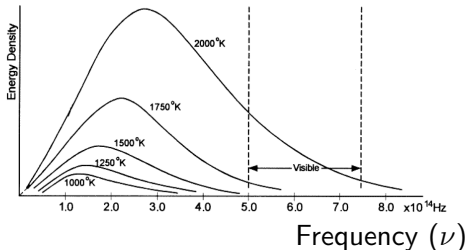
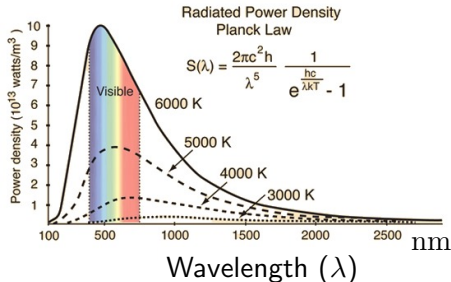
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A black body is an idealized physical body that absorbs all incident **electromagnetic radiation**, regardless of frequency or angle of incidence. In thermal equilibrium (at a constant temperature) it emits electromagnetic radiation called black-body radiation with two notable properties.

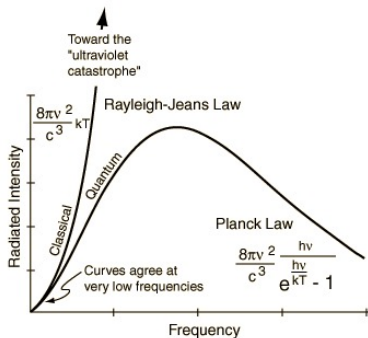
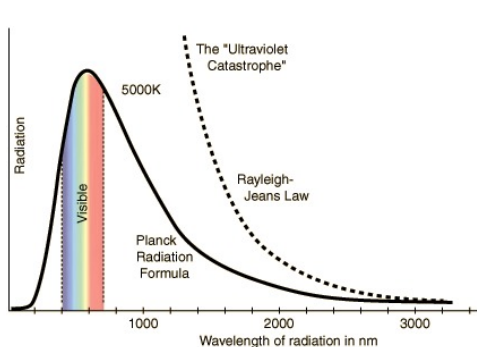
- 1 It is an ideal emitter: it emits as much or more energy at every frequency than any other body at the same temperature.
- 2 It is a diffuse emitter: the energy is radiated isotropically, independent of direction.





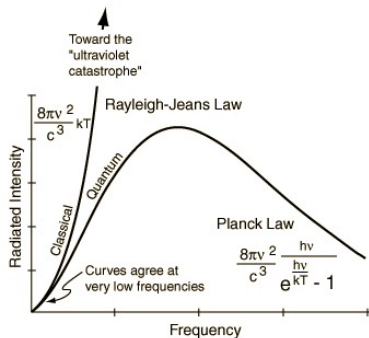
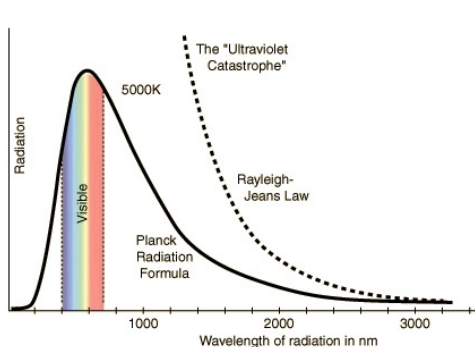
Measured by Lummer and Pringsheim (1899).

$$R_T(\nu)d\nu = \frac{\text{energy}}{\text{time-area}} \quad \text{in the range } \nu \rightarrow \nu + d\nu$$



Rayleigh-Jeans Law

$$u(\nu)d\nu = \frac{8\pi}{c^3} k_B T \nu^2 d\nu \quad \text{in the range } \nu \rightarrow \nu + d\nu$$



$$E = h\nu$$

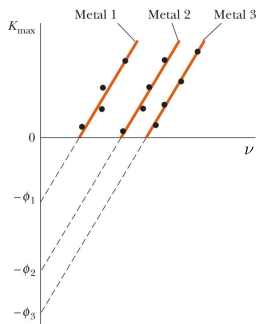
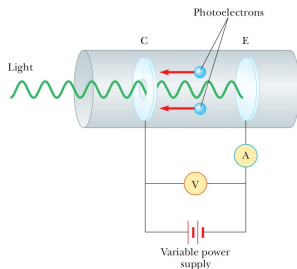
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The Photoelectric Effect - Waves Acting Like Particles 12

- 1 Shine a light on metal and eject electrons.
- 2 Classical physics predicts that any frequency/wavelength of light will work as long as the light is intense enough.
- 3 Measurements by Lennard and others show very different behavior including a linear dependence on frequency and a lower limit. No intensity dependence.
- 4 Einstein uses Planck's hypothesis to explain it with a simple equation invoking the quantum hypothesis

$$K_{max} = eV_{stop} = E - \Phi = h\nu - \Phi$$

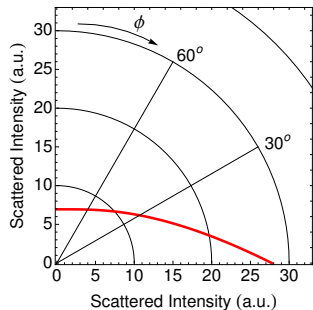
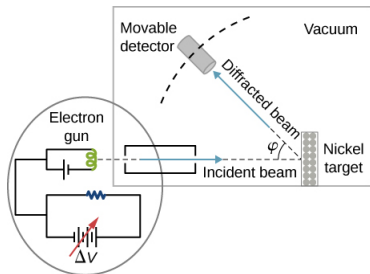
where Φ is the work function, V_{stop} is the minimum voltage for zero current, ν is the frequency of the light, and K_{max} is the maximum kinetic energy of the ejected electrons.



- 1 Shine a beam of electrons on crystalline nickel using the setup here.
- 2 In the polar plot at lower right the scattered electron intensity is proportional to the distance from the origin.
- 3 Classical physics predicts the electrons will collide with the nickel atoms and scatter to large angles. The intensity of scattered electrons should vary smoothly with angle (red curve).
- 4 The data (blue points) disagree wildly.
- 5 de Broglie solves it with electron waves

$$p = \frac{h}{\lambda}$$

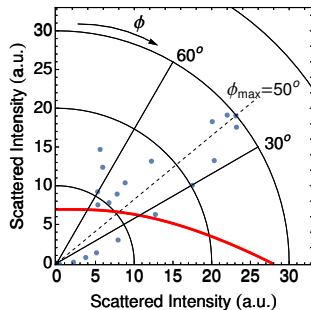
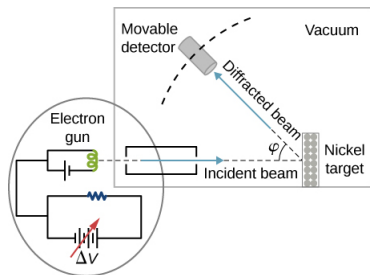
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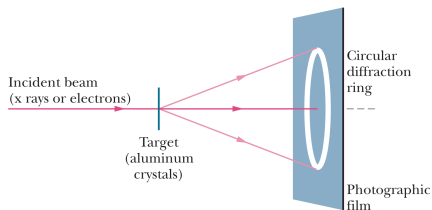


- 1 EM waves transmit energy and momentum.

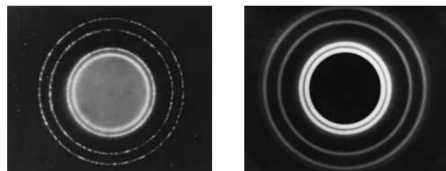
$$p = \frac{\text{energy absorbed}}{c} = \frac{E}{c}$$

- 2 de Broglie invokes symmetry to predict that particles will have wave properties.

$$p = \frac{E}{c} = \frac{h\nu}{c} = \frac{h}{\lambda}$$



Apparatus to test de Broglie hypothesis.



Results for x-rays (left) and electrons (right).

- 1 In classical mechanics there is no limitation on the accuracy of our ability to measure the position $\vec{r}(t)$ and velocity $\vec{v}(t)$ of a particle.
- 2 The only limitations are experimental ones which can be overcome (hopefully) with improvements in technology and technique.
- 3 In wave mechanics (and quantum mechanics) this is no longer true!
- 4 For the motion of a quantum particle in one dimension the Heisenberg Uncertainty Principle is a fundamental limit that cannot be overcome. It is

$$\Delta x \Delta p_x \geq \frac{\hbar}{2}$$

where $\hbar = h/2\pi$, h is Planck's constant and the Δ 's are the uncertainties.

- Spectral lines
- Blackbody radiation
- Photoelectric effect
- Specific heat freeze-out
- Compton effect
- Davisson-Germer
- Radioactivity
- Atomic structure/nuclear and particle physics

The current list:

https://en.wikipedia.org/wiki/List_of_unsolved_problems_in_physics